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CLOUD-FREE LINE-OF-SIGHT CALCULATIONS

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ABSTRACT

A method for computing the probability of having a cloud-free line-of-sight (CFLOS) to or from a given point on the earth, using 3-hourly synoptic weather reports of clouds. The method is based on whole-sky photographs taken during daylight hours over a period of three years at Columbia, Missouri. The computational procedure is an effort to eliminate an apparent oversimplification in previously published data that results from the use of mean cloud cover, an unrelated vertical cloud distribution, and sunshine data. Present results are at variance with those earlier estimates, but compatible with recent observations actually taken from aircraft. Although the CFLOS estimates obtained are, by nature, uncertain, the range of uncertainty was estimated by using published empirical data and a quantitative error analysis.

ACKNOWLEDGMENTS

1 Introduction

When there is an anticipated need to observe a point on the ground from high in the sky (or to observe the sky from a ground station), there may be a distinct advantage in foreknowing the probability of a cloud-free line of sight (CFLOS) for that geographical position. McCabe (1965) and Lund (1965) independently developed semiobjective methods for determining a CFLOS probability, using only climatological data on clouds and sunshine. McCabe's results formed the basis for the calculation of CFLOS probabilities at various stations around the world (Quayle et al., 1968). However, a comparison of these data with data from actual aircraft observations (Bertoni, 1967) suggests that the Quayle estimates of probability were far too high.

McCabe's first assumption is that when bright sunshine is reported at the surface, there is a clear line of sight from the sun to the ground. Relating mean monthly cloud amounts at United States stations to mean number of hours of sunshine, and knowing solar elevation as a function of time, he constructed a graph (Fig. 1) from which could be read the probability of CFLOS as a function of look angle and cloud amount. Then assuming that the basic relationships among these three variables would apply for all cloud heights, McCabe devised a scheme for estimating the distribution of clouds as a function of height, based on the observations of DeBary and Möller (1963), which were made over central Europe.

An outstanding merit of McCabe's work, and also that of Lund (1966), Fig. 2, was that it recognized the important discrepancy between a ground observer's report of cloud amount and the probability of a CFLOS. Also, the concept of using sunshine data as a surrogate for CFLOS is reasonable, although there are discrepancies related to the observational technique that have been universally recognized.
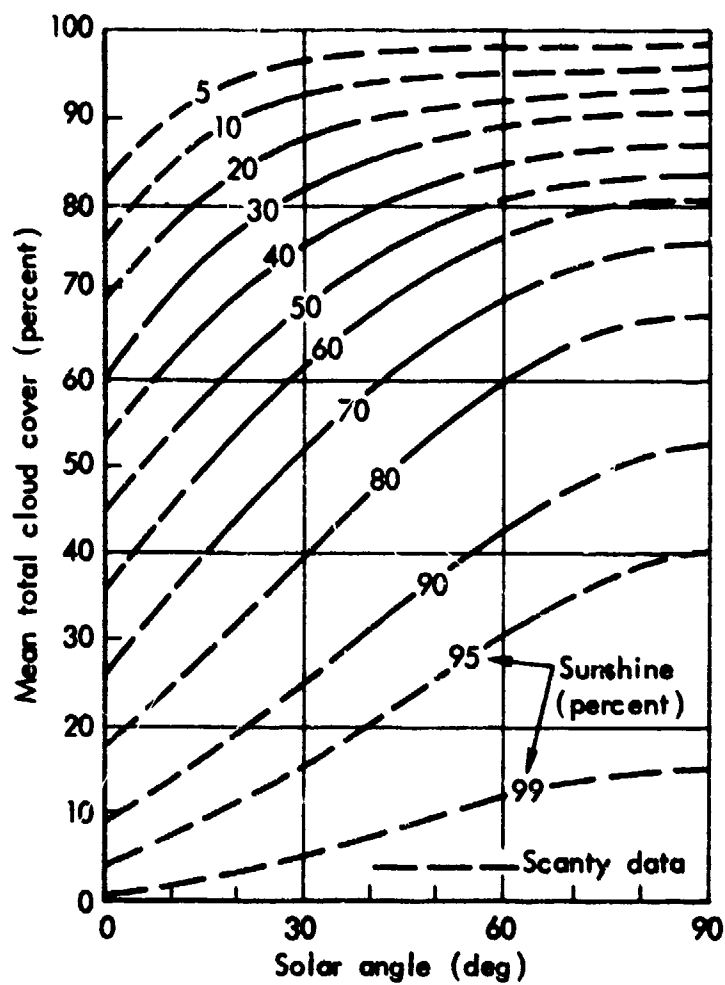
Fig. 1—Approximate mean percentage of
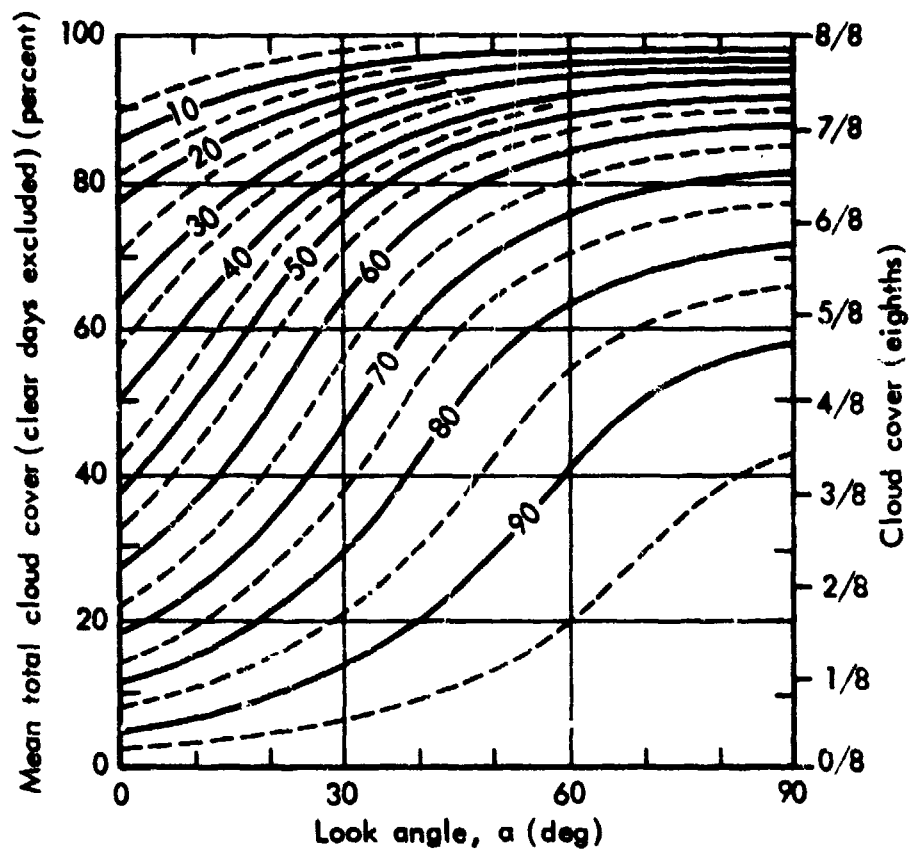sunshine as a function of cloud cover
and sun angle (after McCabe)

Fig. 2 — Probability of a clear line of sight (sunshine) as a
function of percent mean total cloud cover and look
angle a when clear days are excluded from
consideration [after Lund (1966)]

The discrepancy between Quayle's results and Bertoni's observations derives from several effects. The presence of bright sunshine does not have a one-to-one relationship with the ability of a human eye to see an object on the ground. Actually, the CFLOS defined by McCabe neglects all consideration of "thin" clouds or surface haze and refers only to a path through the atmosphere that is free of opaque clouds; so CFLOS probabilities derived by McCabe's method *should* tend to be higher than clear line-of-sight (CLOS) probabilities as reported by Bertoni. Furthermore, studies by Appleman (1962) and Greaves et al. (1971) corroborate that cloudiness reported from above corresponds quite poorly with that reported from below.

When Quayle et al. used the McCabe results to estimate the CFLOS probabilities at many places in the world, they took mean monthly cloud amounts, entered McCabe's graph, and read out probabilities. But the relationship between cloud amounts and CFLOS is nonlinear for all but very shallow look angles, as is shown in Fig. 3 $\left[\text{derived from McCabe (1965)}\right]$, and therefore, averaging the cloud amount before entering the graph can be shown to be poor procedure.

The difficulty is best demonstrated by a *reductio ad absurdum*. Suppose a location had 15 clear days and 15 cloudy days in a month; in this case, both the mean cloudiness and the probability of a CFLOS would be 0.5. If we enter Fig. 1, however, with 0.5 mean cloudiness and a look angle ("solar angle" in Fig. 1) of 90 degrees (either straight up or straight down), we arrive at a CFLOS probability of 0.9.

We decided to perform a critical analysis of the McCabe method without changing the original hypothesis, which has gone a long way toward providing a realistic approach to the CFLOS problem.

We have attempted to reduce some of the uncertainties that have arisen because of certain assumptions in the earlier CFLOS calculations. Specifically, this presentation does the following:

1. Treats a distribution of cloud cover instead of a mean cloud amount at the reporting station.

2. Considers the vertical distribution of the cloud cover using the best available data from an individual weather station.

3. Improves on the estimates made by McCabe and others of relationships among look angle, cloud amount, and CFLCS probability at given ranges.
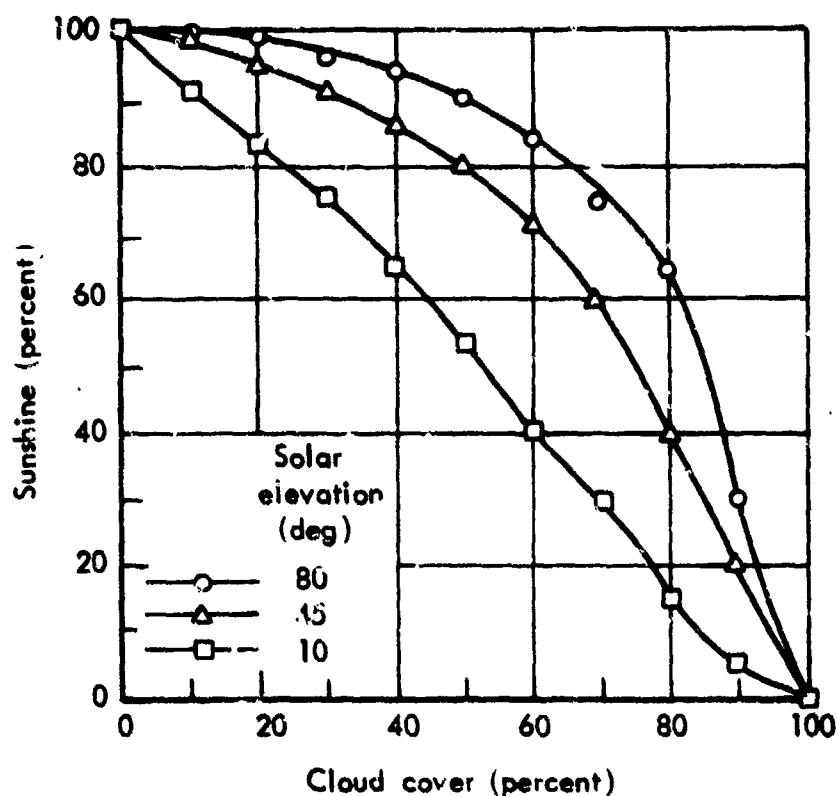
Fig. 3 — Nonlinearity of relationship between CFLOS and cloud cover at three viewing angles (solar elevations)

2 Photogrammetric data

Our methods were applied to data from Columbia, Missouri, where
CFLOS data were visually determined from whole-sky photographs
compiled over a three-year period from March 1, 1966 to February 28,
1969 (Shanklin and Landwehr, 1971).[*] The photographs were taken at
one-hour intervals to reveal the effect of diurnal variations on
cloud cover and to permit correlation of these data with official
Weather Bureau observations taken at the same location. They showed
daytime cloud conditions at nine look angles, 10 through 90 deg,
for the four cardinal compass points starting with true north.
The exact positions of the lines of sight on each print were readily
located with a clear plastic overlay inscribed with 33 small circles
whose centers represent 33 lines of sight at the given azimuths and
look angles. Shanklin and Landwehr present graphs of CFLOS probability
as a function of cloud cover in tenths and look angle for each
azimuth and for all cloud types combined; the graphs indicate the
reliability of the method. For example, the CFLOS probability at
zero cloud cover, which actually represents cloud cover of less
than 5 percent, is greater than 96 percent at all look angles.
Likewise, at ten-tenths clouds, or greater than 95 percent cloud cover,
the probability is less than 9 percent at all look angles.

Shanklin and Landwehr also computed the probabilities of CFLOS
for each azimuth, look angle, and cloud type, on the basis of
sunshine data recorded during the same three-year period by the U.S.
Weather Bureau at Columbia, Missouri, for each tenth of cloud cover.
They found that, for all clouds, these sunshine-based probabilities
of CFLOS varied with sunshine, increasing from an average of 12
percent probability at zero percent sunshine to 71 percent probability
at 100 percent sunshine. These deviations, large when compared with
the above-cited percentages derived from photographs, are due in

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[*] The program was initiated and sponsored by the Air Force
Cambridge Research Laboratories.

part to a characteristic of the Weather Bureau's sunshine recorder:
it does not detect thin clouds. Also, the recorder is directed
toward the position of the sun in the sky (southerly azimuths) only.

For the present study, we averaged the CFLOS probabilities
from the photographs for all azimuths and all clouds (Fig. 4) to
permit direct comparison with McCabe's probabilities (Fig. 1).
In Fig. 4 the curves were slightly smoothed. There were no data
at less than 10 degrees elovation angle, so the character of the
curve from ~ 0 to 10 deg was assumed. Also, photogram data reached
an unexplained maximum of CFLOS probability before 90 deg. This
seems unrealistic and is probably due either to the way the data
were observed, to the effects of lighting at the higher elevation
angles, or to both. Therefore, we flattened the curves for higher
elevation angles, beginning at the point of highest CFLOS. This
point generally occurred at elevation angles greater than 60 deg.
A more complete discussion of the data of Shanklin and Landwehr
is given in Lund and Shanklin (1972).

We have used the CFLOS probability data in Fig. 4 for our
calculations largely because these probabilities are based on
carefully checked measurements of low, middle, and high clouds.
When these probabilities are averaged for all clouds and all
azimuths, they fall generally within 10 percent of the McCabe
estimates, as is shown in Table 1. It should be noted that the
averages from the photographs give a CFLOS probability up to 10
percent lower than that for sunshine data (Fig. 1) for cloud
amounts equal to or less than 6/8 (0.8) cloud cover, but up to
8 percent higher for greater cloud amounts, an important factor,
particularly when considering a cloudy area such as Columbia,
Missouri or, the eastern United States.
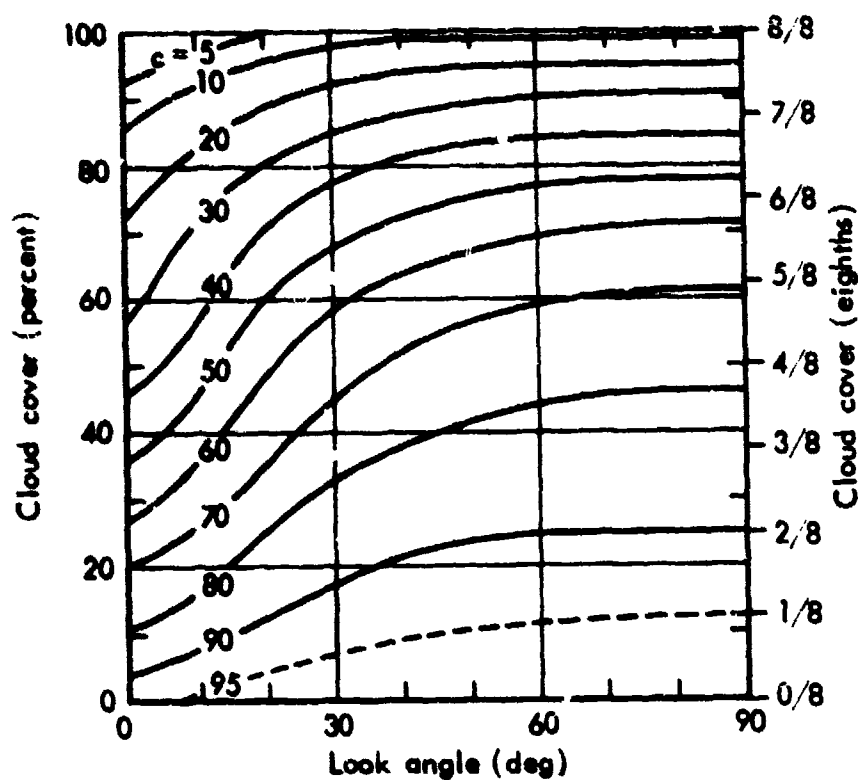
Fig. 4 — Probability (c) of CFLOS: All azimuths, all cloud covers (from Shanklin and Landwehr, Table 9)

Table 1

COMPARISON OF CFLOS PROBABILITIES FROM TWO SOURCES

Sources of Data and (Difference)[a]	Cloud Cover, k								
	8/8	7/8	6/8	5/8	4/8	3/8	2/8	1/8	0/8
	Elevation Angle 30 deg								
Shanklin/Landwehr	0.06	0.27	0.4.	0.56	0.68	0.76	0.85	0.92	0.99
McCabe	0.00	0.20	0...	0.58	0.72	0.82	0.90	0.96	1.00
(Difference)	+0.06	+0.07	+0.03	-0.02	-0.04	-0.06	-0.05	-0.04	-0.01
	Elevation Angle 45 deg								
Shanklin/Landwehr	0.07	0.32	0.49	0.61	0.72	0.81	0.89	0.94	0.99
McCabe	0.00	0.28	0.51	0.70	0.80	0.88	0.94	0.98	1.00
(Difference)	+0.07	+0.04	-0.02	-0.09	-0.08	-0.07	-0.05	-0.04	-0.01
	Elevation Angle 60 deg								
Shanklin/Landwehr	0.08	0.35	0.52	0.67	0.76	0.83	0.90	0.95	0.99
McCabe	0.00	0.33	0.62	0.77	0.85	0.93	0.96	0.99	1.00
(Difference)	+0.08	+0.02	-0.10	-0.10	-0.09	-0.10	-0.06	-0.04	-0.01

[a] Shanklin/Landwehr estimate minus McCabe estimate.

3   Synoptic data

Data on the frequency of various amounts of cloud cover
(in eighths[*]) for each cloud height considered in this report
(Table 2) were obtained from magnetic tapes in the Air Force
TDF-14 format, prepared by the Environmental Technical Applications
Center (ETAC). This format combines aviation reports with supple-
mental cloud data, every three hours. The data are available for
all regularly reporting stations across the United States.
(Overseas, the Air Force TDF-13 format is used.)

These tapes record synoptic data normally appearing on circuit
"C". The program will accept records of any period greater than one
year; however, the Columbia data covered a span from January 1, 1945
to December 1, 1968. The taped data were placed in the Rand Weather
Data Bank (RAWDAB). They were then interrogated through a special
set of computer programs developed by R. E. Huschke and E. Rodriguez
for the cloud amount and height. The height listing eliminated the
need for the DeBary and Müller (1963) vertical distribution of clouds.

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[*]The choice of eighths of cloud as a breakdown was one of
expedience, as many of the world's cloud data are reported in
eighths. Lund (1965) achieved rather good results by merely
separating clear days from days with cloud and using the mean cloud
amount for the latter. It may therefore be reasonably assumed that
a finer breakdown is unwarranted. A coarser breakdown, on the other
hand, would necessitate additional manipulation of the cloud data
before entering the calculation.

Table 2

CFLOS PROBABILITIES FOR WINTER AT COLUMBIA, MISSOURI, BASED ON EIGHTHS OF CLOUD COVER AT LEVELS GIVEN

(Data from daylight hours, based on 11,804 reports)

Level	Height (ft)	e/8 [b]	\ Cloud-cover probabilities at, and below, height given\ 8/8	7/8	6/8	5/8	4/8	3/8	2/8	1/8	0/8	CFLOS[a] for three look angles at each height\ 10°	45°	60°
30	30,000	0.0	0.6665	0.0481	0.0660	0.0220	0.0114	0.0241	0.0585	0.0356	0.0675	0.2692	0.2841	0.3016
29	25,000	0.0	0.6636	0.0467	0.0636	0.0212	0.0112	0.0233	0.0565	0.0342	0.0798	0.2754	0.2938	0.3068
28	20,000	0.0	0.6586	0.0445	0.0605	0.0232	0.0102	0.0222	0.0537	0.0324	0.0970	0.2846	0.3024	0.3151
27	19,000	0.0	0.6503	0.0418	0.0570	0.0191	0.0106	0.0210	0.0505	0.0303	0.1192	0.2978	0.3147	0.3270
26	18,000	0.0	0.6414	0.0391	0.0535	0.0180	0.0102	0.0197	0.0472	0.0283	0.1426	0.3114	0.3277	0.3395
25	17,000	0.0	0.6317	0.0364	0.0500	0.0169	0.0098	0.0185	0.0440	0.0262	0.1665	0.3254	0.3414	0.3527
24	16,000	0.0	0.6214	0.0336	0.0465	0.0159	0.0094	0.0173	0.0407	0.0242	0.1910	0.3493	0.3554	0.3666
23	15,000	0.0	0.6104	0.0309	0.0431	0.0149	0.0090	0.0161	0.0376	0.0223	0.2158	0.3553	0.3704	0.3809
22	14,000	0.0	0.5989	0.0282	0.0398	0.0139	0.0087	0.0150	0.0345	0.0205	0.2405	0.3721	0.3854	0.3955
21	13,000	0.0	0.5871	0.0256	0.0367	0.0130	0.0083	0.0140	0.0317	0.0188	0.2649	0.3879	0.4006	0.4102
20	12,000	0.0	0.5751	0.0232	0.0338	0.0123	0.0080	0.0130	0.0291	0.0172	0.2883	0.4035	0.4155	0.4248
19	11,000	0.3	0.5632	0.0211	0.0312	0.0116	0.0076	0.0122	0.0267	0.0159	0.3104	0.4196	0.4300	0.4389
18	10,000	0.0	0.5518	0.0192	0.0290	0.0110	0.0074	0.0115	0.0247	0.0148	0.3306	0.4326	0.4435	0.4521
17	9,000	0.0	0.5414	0.0176	0.0272	0.0105	0.0071	0.0109	0.0231	0.0139	0.3484	0.4452	0.4557	0.4640
16	8,000	0.0	0.5322	0.0163	0.0257	0.0102	0.0069	0.0105	0.0218	0.0132	0.3632	0.4561	0.4667	0.4742
15	7,000	0.0	0.5247	0.0153	0.0247	0.0099	0.0068	0.0102	0.0209	0.0123	0.3748	0.4647	0.4746	0.4824
14	6,000	0.0	0.5051	0.0143	0.0228	0.0091	0.0064	0.0095	0.0195	0.0121	0.4015	0.4858	0.4951	0.5078
13	5,000	0.0	0.4861	0.0122	0.0206	0.0082	0.0058	0.0086	0.0174	0.0110	0.4361	0.5126	0.5213	0.5282
12	4,000	0.0	0.4480	0.0099	0.0180	0.0072	0.0049	0.0071	0.0144	0.0092	0.4813	0.5472	0.5540	0.5613
11	3,000	0.3	0.4063	0.0072	0.0150	0.0060	0.0045	0.0052	0.0104	0.0065	0.5400	0.5919	0.5984	0.6030
10	2,000	0.0	0.3521	0.0044	0.0115	0.0046	0.0019	0.0030	0.0059	0.0033	0.6133	0.6436	0.6538	0.6583
9	1,500	0.3	0.2847	0.0022	0.0081	0.0033	0.0007	0.0013	0.0026	0.0012	0.6958	0.7167	0.7205	0.7239
8	1,000	0.3	0.2140	0.0008	0.0051	0.0021	0.0002	0.0003	0.0007	0.0002	0.7764	0.7864	0.7891	0.7915
7	700	0.0	0.1486	0.0002	0.0029	0.0013	0.0000	0.0001	0.0001	0.0000	0.8467	0.8494	0.8512	0.8528
6	500	0.0	0.0957	0.0001	0.0016	0.0007	0.0003	0.0000	0.0000	0.0000	0.9019	0.8997	0.9008	0.9019
5	400	0.0	0.0575	0.0000	0.0008	0.0004	0.0000	0.0000	0.0000	0.0000	0.9412	0.9359	0.9365	0.9371
4	300	0.0	0.0287	0.0000	0.0003	0.0002	-0.0000	0.0000	0.0000	-0.0000	0.9708	0.9630	0.9634	0.9637
3	200	0.0	0.0101	0.0000	0.0001	0.0000	-0.0000	-0.0000	-0.0000	-0.0000	0.9897	0.9805	0.9806	0.9807
2	100	0.0	0.0015	-0.0003	0.0000	0.0000	-0.0000	-0.0000	-0.0000	-0.0000	0.9984	0.9986	0.9986	0.9886
1	00	0.0	-0.0000	-0.0005	-0.0000	-0.0000	-0.0000	-0.0000	-0.0000	-0.0000	1.0000	0.9900	0.9900	0.9900

[a] CFLOS based on Shanklin and Landwehr

[b] Obscured

4 Method of calculation

In order to reduce significantly the inaccuracies that result from using mean cloud amount values, because of the nonlinearity, we established the overall probability of a CFLOS for each reported cloud amount separately and then computed an average. To do this required (1) the RAWDAB frequency distribution of clouds to satisfy the function $f(\ell, k)$, where ℓ is the height of the cloud and k is the amount in eighths, and (2) the photographically derived probability of a CFLOS as a function of look angle, α, and cloud amount, k, represented by the function $c(\alpha, k)$. Now, if $f(\ell, k)$ represents the frequency of clouds of k eighths at height ℓ, and $c(\alpha, k)$ represents the probability of a CFLOS at a look angle α through clouds in the amount k, then the expected probability of a CFLOS from the surface to a given height (or from a given height to the surface) at a given look angle is the sum of the frequency, f, multiplied by the probability, c:

$$P(\alpha, \ell) = \sum_{k=0}^{8} c(\alpha, k)f(\ell, k) \tag{1}$$

The next task was to examine critically the functions c and f. In McCabe's original work, the value of $c(\alpha, k)$ was estimated from mean monthly cloud amounts and mean monthly sunshine records. Lund (1965) gave a thorough discussion of the entire problem, which details some of the subtleties of the CFLOS problem, and in a later paper (Lund, 1966), he presented a graph similar to McCabe's but based on the average cloudiness of only those days that were not completely clear. Lund's graph of the function labeled c in this report is shown in Fig. 2. As is stated earlier, the aircraft observations reported by Bertoni (1967) were useful in calling into question the use of sunshine data, but since they were not related in a one-to-one fashion with ground observations, they could not be used effectively

in constructing the function we desire. Shanklin and Landwehr (1971), however, present a wealth of data on the variations of CFLOS for all clouds as a function of look angle and cloud amount. Therefore in our judgement, the function $c(\alpha, k)$, as shown in Fig. 4, represents the best estimate available at this time.

If the function presented in Fig. 4 is the probability of seeing the ground from any altitude, as long as k is interpreted as the cloud amount *below* the altitude in question, then the function $f(\ell, k)$ should represent the frequency that there are k eighths of cloud below level ℓ. The questionable validity of the DeBary and Möller distribution (used by McCabe) for areas other than central Europe has caused some concern over the use of the method. In order to eliminate this objection, the observed vertical and horizontal distributions of clouds were extracted from the same magnetic tapes of daily synoptic reports made available by ETAC. The difficulty that arises with this approach is that it is necessary to rely on the estimates of height made by ground observers. Estimates of cloud height and amount made by ground observers are deficient in three ways: (1) low clouds obscure the extent of higher cloud coverage, (2) the cloud-height reporting code does not provide a consistent scale, and (3) observers tend to have biases in the heights that they do report. Historical weather records, therefore, do not give smooth distributions of cloud amount with height. The original $f(\ell, k)$ function for levels to 30,000 feet exhibited clumping tendencies which were carried through to the computation of the CFLOS, thus generating irregular curves. Since there is no a priori reason to expect the true cloud distribution to be so discontinuous, we decided to apply a smoothing technique to the vertical distribution function. We found a log-normal distribution that nicely fit the low clouds and another log-normal that could fit the middle and high clouds. The final result is an equation, for each eighth of cloud cover, of the form

$$f(\ell, k) = \begin{cases} a \, \ell n \, N(m_1, \, \sigma_1) & \text{for } 1 \le \ell \le 14 \\ \\ b \, \ell n \, L(m_2, \, \sigma_2) & \text{for } 15 \le \ell \le 30 \end{cases} \tag{2}$$

where $\ell n \, N(m, \sigma)$ represents the normal distribution in the logarithm
of height; the subscripts 1 and 2 refer to the lower ($< 6,000$ feet)
and upper clouds (6,000 to 30,000 feet), respectively (Table 2);
and a and b are weighting factors depending on the relative amounts
of lower and upper clouds. Figure 5 shows the observed data for 8/8
cloud cover over Columbia in the winter. The peaks reported for
middle and high clouds are obvious, as is the discontinuous nature
of the distribution of the low clouds. Also shown in Fig. 5 are the
two sections of the component log-normal fit to the data. We believe
that the smoothing achieved by the fitting procedure is probably a
more realistic representation of cloud occurrence than is the obviously
clumped data as reported. The mathematical details of the smoothing
procedure are as follows:

$$x_\ell = \text{height interval; } x_\ell = \ell$$

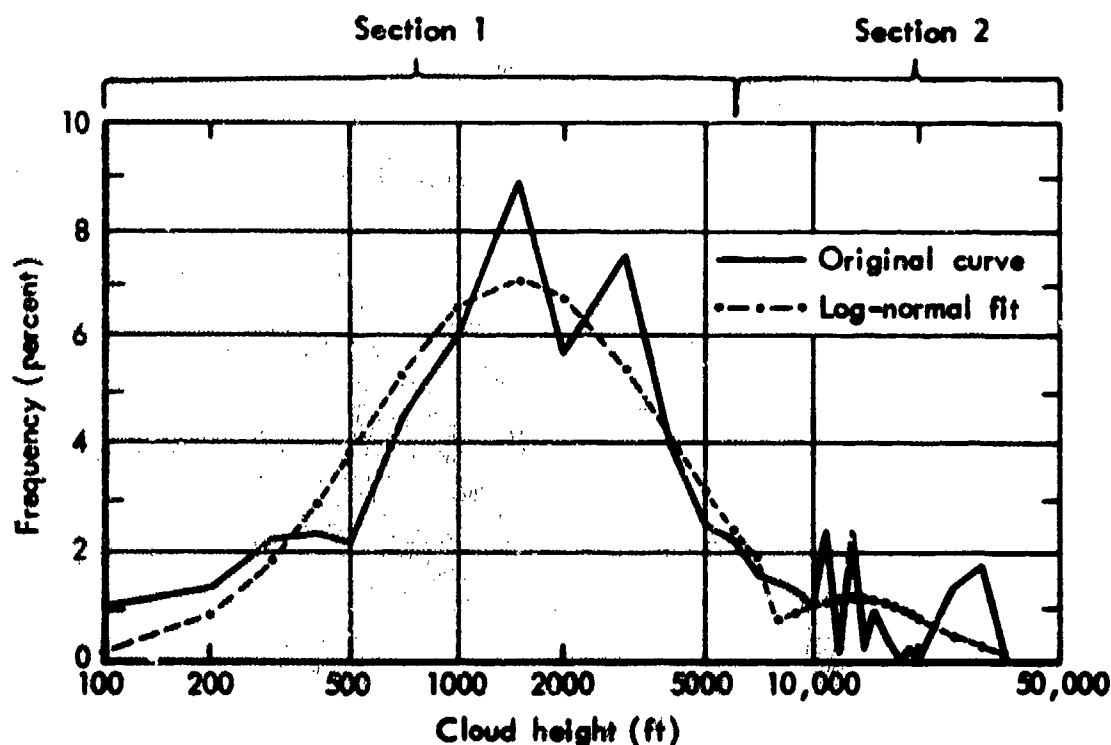$$g_\ell = \text{frequency of cloud at height interval } \ell$$

Fig.5 — Cloud-cover frequency at various altitudes over Columbia,
Missouri (1945-1968), and log-normal curves used for
smoothing cloud-height data

g_{1_ℓ} = normalized g_ℓ with respect to intervals from 0 to 6,000 ft

$$= \frac{g_\ell}{\sum_{\ell=1}^{14} g_\ell}, \qquad \ell = 1, 2, \ldots, 14$$

g_{2_ℓ} = normalized g_ℓ with respect to intervals from 6,000 to 30,000 ft

$$= \frac{g_\ell}{\sum_{\ell=14}^{30} g_\ell}, \qquad \ell = 14, 15, \ldots, 30$$

$\bar{x}_1$ = mean height for section 1 = $\sum_{\ell=1}^{14} g_{1_\ell} \ell n \, x_\ell$

$\bar{x}_2$ = mean height for section 2 = $\sum_{\ell=14}^{30} g_{2_\ell} \ell n \, x_\ell$

σ_1 = standard deviation for section 1

$$= \left[\sum_{\ell=1}^{14} g_{1_\ell} (\ell n \, x_\ell - \bar{x}_1)^2 \right]^{1/2}$$

σ_2 = standard deviation for section 2

$$= \left[\sum_{\ell=14}^{30} g_{2_\ell} (\ell n \, x_\ell - \bar{x}_2)^2 \right]^{1/2}$$

Note that if all of the data reported in section i (1 or 2) fall within one interval of height, then σ_1 = 0, and the approximation breaks down. Therefore, we followed the convention that σ_1 = 0.25 in such cases.

Let

$$
h(\ell, k) = \begin{cases} \dfrac{\alpha_1}{\sigma_1 \sqrt{2\pi}} \ \exp\left[-\dfrac{(\ell n \, x_\ell - \bar{x}_1)^2}{2\sigma_1^2} \right] & \text{for} \quad 1 \leq \ell \leq 14 \\[4em] \dfrac{\alpha_2}{\sigma_2 \sqrt{2\pi}} \ \exp\left[-\dfrac{(\ell n \, x_\ell - \bar{x}_2)^2}{2\sigma_2^2} \right] & \text{for} \quad 15 \leq \ell \leq 30 \end{cases}
$$

where α_1 and α_2 are normalizing factors for sections 1 and 2 respectively. Thus the approximation to the curve (0 to 30,000 ft) is the following combination of log-normal distributions:

$$
f'(\ell, k) = \begin{cases} a \ \cdot \ h(\ell, k) & \text{for} \quad 1 \leq \ell \leq 14 \\[1em] b \ \cdot \ h(\ell, k) & \text{for} \quad 15 \leq \ell \leq 30 \end{cases}
$$

where

$$
a = \frac{\displaystyle\sum_{\ell=1}^{14} s_\ell}{\displaystyle\sum_{\ell=1}^{30} s_\ell + s_{14}}
$$

and

$$
b = \frac{\displaystyle\sum_{\ell=14}^{30} s_\ell}{\displaystyle\sum_{\ell=1}^{30} s_\ell + s_{14}}
$$

The approximation and the original curve are compared in Fig. 5.

The distribution function $f(\ell, k)$ was smoothed by using the above method for each fraction (in eighths) of cloud cover (1/8 to 8/8). A cumulative $D(\ell, k)$ table (Table 2) was then generated from the smoothed $f(\ell, k)$ table. Table 2 represents the probability of finding k eighths of cloud cover at and below the height interval ℓ. The 0/8 entry was computed as

$$1 - \sum_{k=1}^{8} D(\ell, k), \quad \text{for } \ell = 1, 2, \ldots, 30$$

The entries in the cumulative table are used to compute the CFLOS at a given look angle α, and *below* height interval ℓ (Table 2, three right-hand columns):

$$\text{CFLOS } (\alpha, \ell) = \sum_{k=0}^{8} c(\alpha, k)D(\ell, k)$$

where $c(\alpha, k)$ are CFLOS probabilities based on Shanklin and Landwehr data.

Our computational procedure for any station and for any season is, then, as follows:

1. Extract the observed vertical distribution of clouds for the station and season for each eighth of cloud cover from the RAWDAB tapes.

2. Apply the smoothing technique of Eq. (2) to obtain a smooth distribution with height for each eighth of cloud cover.

3. Construct a cumulative distribution for each eighth of cloud cover (such a distribution is illustrated in Table 2).

4. Obtain from the distribution the values of $c(\alpha, k)$ for all desired look angles for each value of k, i.e., for each cloud cover amount, in eighths.

5. For each elevation, perform the multiplication and summation indicated by Eq. (1) for each desired look angle.

6. Transform the CFLOS probability as a function of height and look angle to a probability as a function of range and look angle by substit ℓ R = ℓ/sin α for height in the probability function (see Ta   3).

Some illustrative results are shown in the three right-hand columns of Table 2. Although this procedure sounds complicated, it can be quickly and easily carried out by a set of computer programs. For this study, we considered only Columbia, Missouri, during four seasons. Look angles of 30, 45, and 60 deg and thirty elevations (up to 30,000 ft) were examined.

Table 3

HEIGHT/RANGE CONVERSION TABLE

| Height | | Range, R (km) | | | Height | | Range, R (km) | | |
| | | Look Angle (deg) | | | | | Look Angle (deg) | | |
ft	km	60	45	30	ft	km	60	45	30
100	0.03	0.04	0.04	0.06	11,000	3.35	3.87	4.74	6.70
200	0.06	0.07	0.09	0.12	12,000	3.66	4.23	5.18	7.32
300	0.09	0.10	0.13	0.18	13,000	3.96	4.57	5.60	7.92
400	0.12	0.14	0.17	0.24	13,124	4.00	4.62	5.66	8.00
500	0.15	0.17	0.21	0.30	14,000	4.27	4.93	6.04	8.54
700	0.21	0.24	0.30	0.42	15,000	4.57	5.28	6.46	9.14
1,000	0.31	0.36	0.44	0.62	16,000	4.88	5.64	6.90	9.76
1,500	0.46	0.53	0.65	0.92	16,405	5.00	5.77	7.07	10.00
2,000	0.61	0.70	0.86	1.22	17,000	5.18	5.98	7.33	10.36
3,000	0.91	1.05	1.29	1.82	18,000	5.49	6.34	7.76	10.98
3,281	1.00	1.16	1.41	2.00	19,000	5.79	6.69	8.19	11.58
4,000	1.21	1.40	1.71	2.42	19,686	6.00	6.93	8.49	12.00
5,000	1.52	1.76	2.15	3.04	20,000	6.10	7.04	8.63	12.20
6,000	1.83	2.11	2.59	3.66	22,967	7.00	8.08	9.90	14.00
6,562	2.00	2.31	2.83	4.00	25,000	7.62	8.80	10.78	15.64
7,000	2.13	2.46	3.01	4.26	26,248	8.00	9.24	11.31	16.00
8,000	2.44	2.82	3.45	4.88	29,529	9.00	10.39	12.73	18.00
9,000	2.74	3.16	3.88	5.48	30,000	9.14	10.55	12.93	18.28
9,843	3.00	3.46	4.24	6.00	32,808	10.00	11.54	14.14	20.00
10,000	3.05	3.52	4.31	6.10					

5   Error analysis

The root-mean-square errors of five different methods of estimating CFLOS, together with their biases, were computed by Lund (1966). Lund's criterion for estimating errors was sunshine data--a portion of the same data from which he derived the coefficients for the five methods. Unfortunately, we have no such objective measures against which our computed probabilities can be tested. Therefore, for our error analysis, it was necessary to estimate the magnitude of the errors in the c and f functions and propagate them through the computation indicated in Eq. (1). The error-propagation rules, assuming independence of errors, are given in Worthing and Geffner (1943), i.e.,

$$p^2(\ell, \alpha, k) = f^2(\ell, k)p_c^2 + c^2(\alpha, k)p_f^2 \qquad (3)$$

$$p(\ell, \alpha) = \left[\sum_{k=0}^{8} p^2(\ell, \alpha, k)\right]^{1/2} \qquad (4)$$

where p_c is the probable error in c; p_f is the probable error in f; and $p(\ell, \alpha)$ is the probable error in the CFLOS estimate.

We estimated p_c by first comparing the smoothed curves of Fig. 4 with the raw data of Shanklin and Landwehr (1971) for a look angle of 45 deg. The discrepancies ranged from -1.0 percent to +0.3 percent. Comparing the smoothing done on the same data by two analysts, we found that the maximum difference was 1.0 percent. For the purposes of this demonstration, therefore, we assume a probable error of 1.0 percent as a safe estimate for the function $c(\alpha, k)$.

The probable errors of the function $f(\ell, k)$ are a bit harder to derive. The standard errors in estimates of the probability of occurrence of a cloud amount for the raw data range from less than 1.0 percent (for the more frequent occurrences) to 5.0 percent (for the less frequent occurrences). The effect of the smoothing is difficult to estimate, but for purposes of crudely estimating

the reliability, we will assume that p_f (the probable error of the function f) is 3.0 percent.

Using the above estimates of the probable errors and using values of c taken from Fig. 4 for a look angle of 45 deg, and values of f taken from Table 2 for cloud heights of 2,000, 4,000, and 6,000 ft, we evaluated Eqs. (3) and (4). In all cases, the first term on the right-hand side of Eq. (3) was so much less than the second term that the variation with altitude was undetectable, because c and p_f are assumed to be invariant with height. The resultant probable error at a 45-deg look angle was found to be $\pm$ 6.5 percent--a value that is consonant with the results of Lund. These errors are plotted in Fig. 6 for comparison with the results based on the data of Quayle et al. This comparison strongly suggests that those estimates of CFLOS probability are entirely too high.
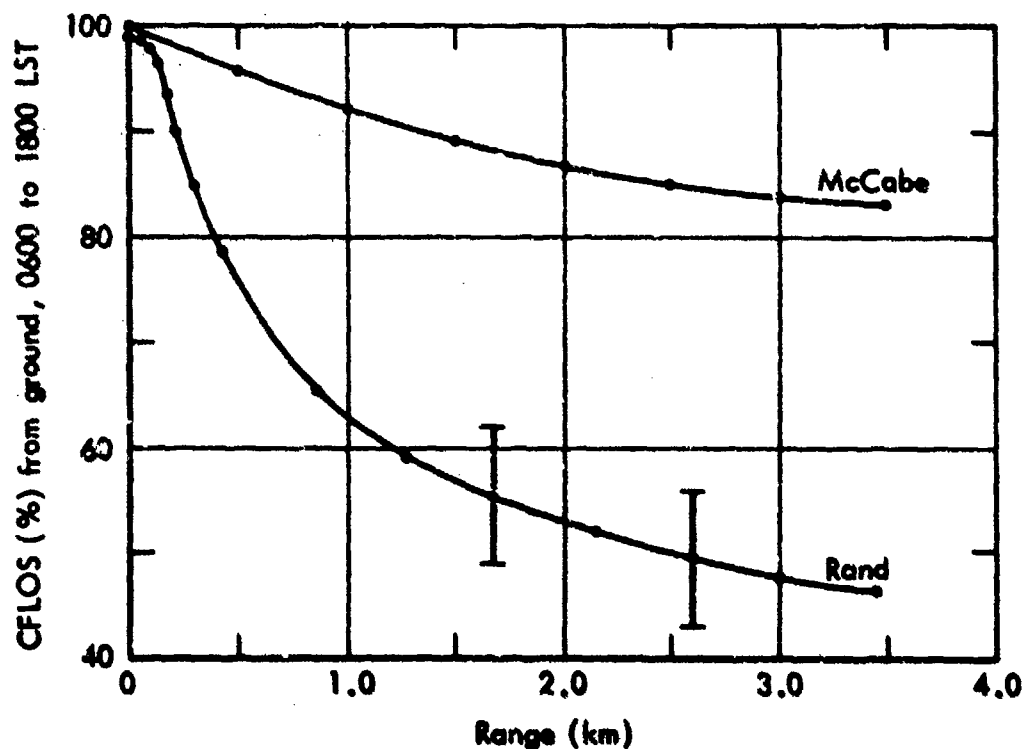

Fig. 6—CFLOS at 45-degree look angle over Columbia, Missouri, (winter) compared with McCabe's CFLOS results

6   Seasonal CFLOS probabilities at Columbia, Missouri

An example of the patterns of CFLOS probabilities, looking
upward from the ground during the hours 0600 — 1800 LST (local
standard time) and using the techniques just discussed, is shown
for Columbia, Missouri (38°58'N-92°22'W) (Figs. 7 and 8). The
seasons are paired so that summer (June-July-August) and winter
(December-January-February) can be compared on one graph, and
spring (March-April-May) and fall (September-October-November)
on another. The curves indicate the probability of CFLOS for
three viewing angles at short ranges (left) and long ranges
(right). They were plotted with the FR-80 Integrated Graphics
System.

Columbia (and much of the eastern United States) is dominated
by unstable tropical maritime (mT) air in summer, and modified,
but stable, polar continental (cP) air in winter. Winter data from
ETAC from 1945 through 1960 show that Columbia is clear on ~8
percent of the days and has scattered (<5/8) clouds on only 10
percent. The remaining broken-to-overcast cloud cover results
from modification along the flow or by interaction of cP and moist
mT along frontal lines. Resulting low stratus and stratocumulus-type
clouds dominate and give the rapid dip of the winter curves
(Fig. 7) at the short ranges. The "bump" in the curves between
4 and 6 km indicate the presence of middle clouds, largely from
frontal overrunning during this season.

Summer weather is less complicated by extensive storms,
although line squalls and thunderstorms with multiple layers account
for much of the erratic slope in these curves (Fig. 7). The predominant
clouds are cumulus, within the dominating moist, unstable mT air mass.
They form at a higher level than the winter stratus and thereby
allow a more gradual decrease in the CFLOS.

The fall and spring curves (Fig. 8) show similar characteristics.
Since there is a rapid decrease in the CFLOS probability at the short
ranges, and a marked "bump" at middle ranges, these curves indicate
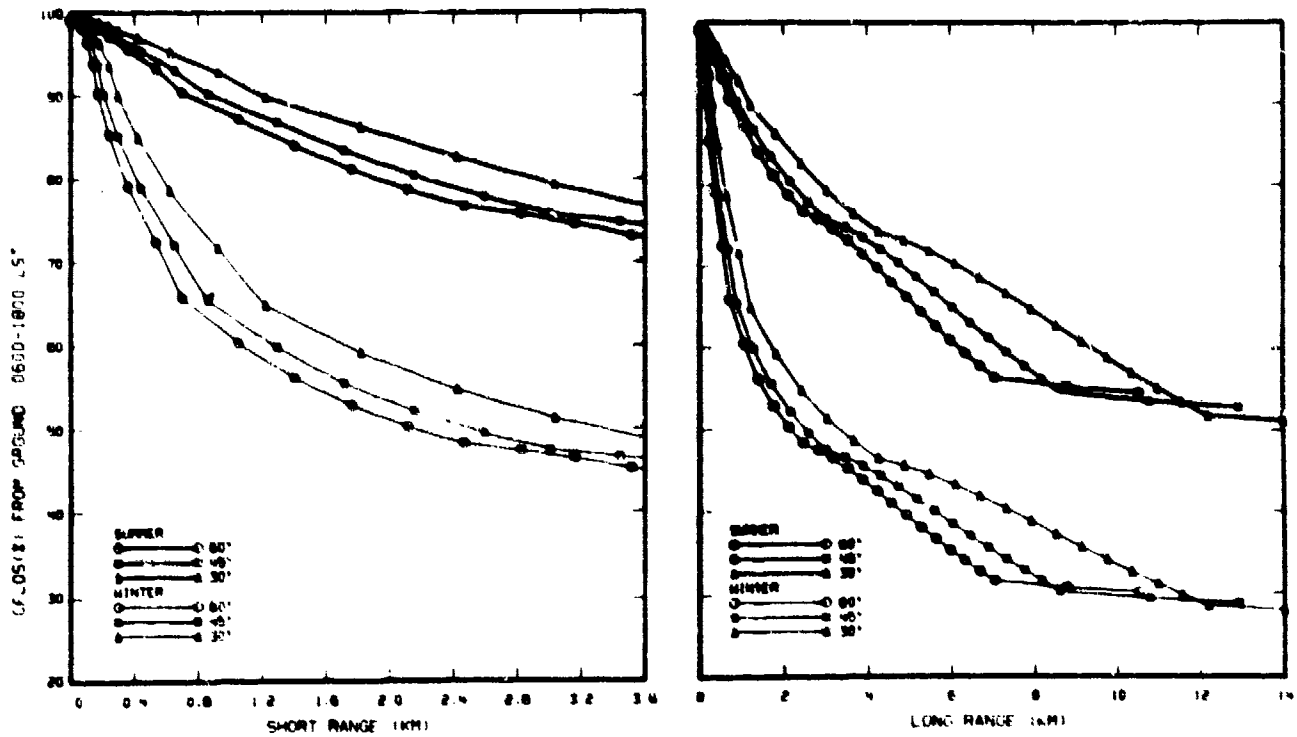a predominant winter influence.

Fig. 7 – CFLOS probabilities, summer and winter, at Columbia, Missouri
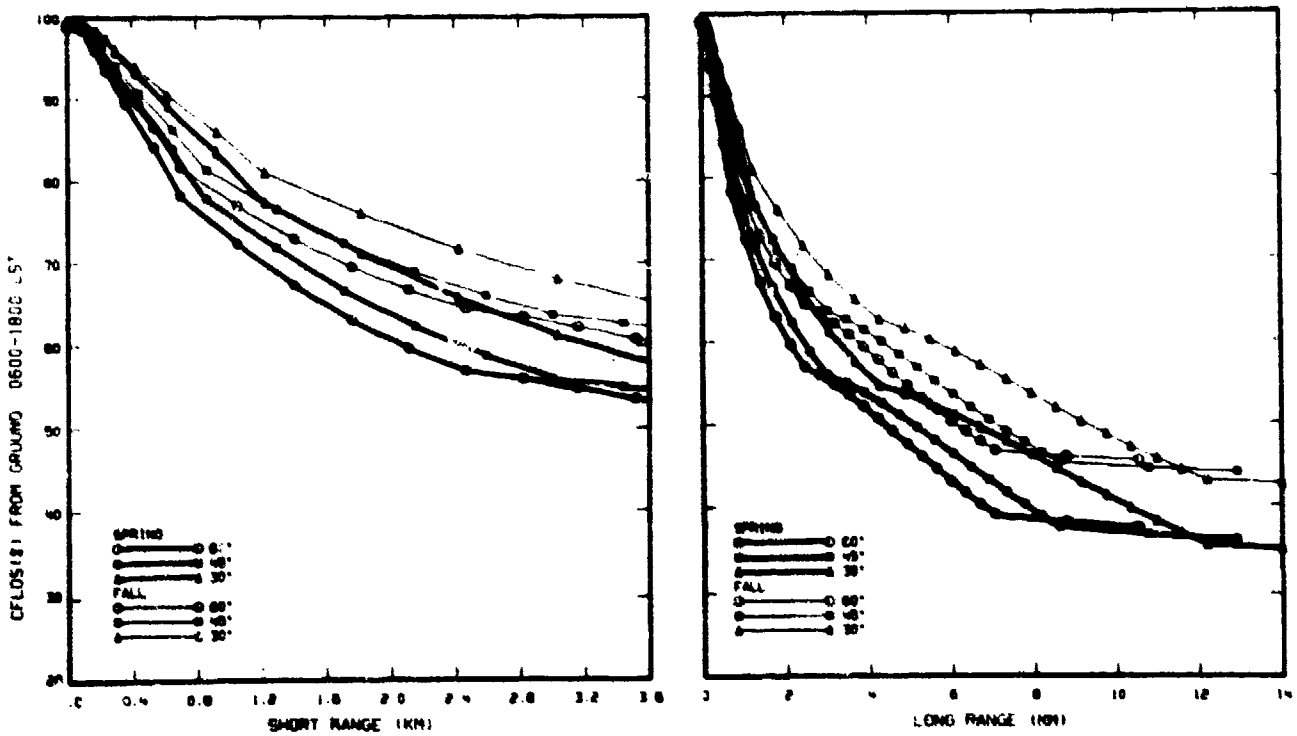

Fig. 8 – CFLOS probabilities, spring and fall, at Columbia, Missouri

7   Concluding remarks

The use of station data for determining eighths of cloud cover
at given heights, and the use of CFLOS measurements for known clouds,
provide greater flexibility for determining local CFLOS probabilities
around the world at given *ranges*. These CFLOS data are clearly more
sensitive to the climate zone of the stations of interest than are
earlier data.

As is stated at the beginning of this report, we reviewed the
entire computational procedure for estimating CFLOS probabilities
originally developed by McCabe (1965) because estimates based on that
work in Quayle et al. (1968) seemed unreasonably optimistic. The
CFLOS probabilities for a 45-deg look angle estimated by Quayle et al.,
for Columbia in the winter, were compared with our results in Fig. 6.
We have carefully calculated this smooth curve and believe that the
detail we have captured at the short ranges is realistic. We believe
it properly reflects the asymptotic behavior at longer ranges and
that the differences in probabilities between, say, the 50 percent
and the 85 percent shown here for 2.6 km range excludes any credence
that the Quayle curve truly represents the probability of a CFLOS.

From the information currently available, it appears that
Shanklin and Landwehr's measurements in Fig. 4 for $c(\alpha, k)$ are probably
the best data presently available. These are *measurements* of actual
cloud conditions, whereas McCabe's data represent *estimates* of the
difference between ground observations and what might be seen from
above. Refinement of Fig. 4 will require a continued program of
direct measurement at Columbia, Missouri, and elsewhere to establish
a good climatological base, especially as related to the effects of
different cloud genera on CFLOS.

Measurements made without reference to surface observation, such
as those of Bertoni, provide some insight into the CFLOS problem and
have been helpful, but they do not provide a realistic means for utilizing
available climatological data.

Findings from direct-measurement programs such as Appleman's and that of Shanklin and Landwehr might constitute the final steps for establishing definitive CFLOS probabilities. Huschke (1971) outlines a scheme for relating these data to air masses of similar origins and to the general circulation patterns of both the northern and southern hemispheres. Therefore, data collected according to this scheme at carefully chosen test sites across the United States could form the basis for relationships between ground observations and CFLOS that would be climatologically reliable for worldwide application, especially in areas where data are sparse.

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